



JET SUPPRESSION AND ENERGY LOSS

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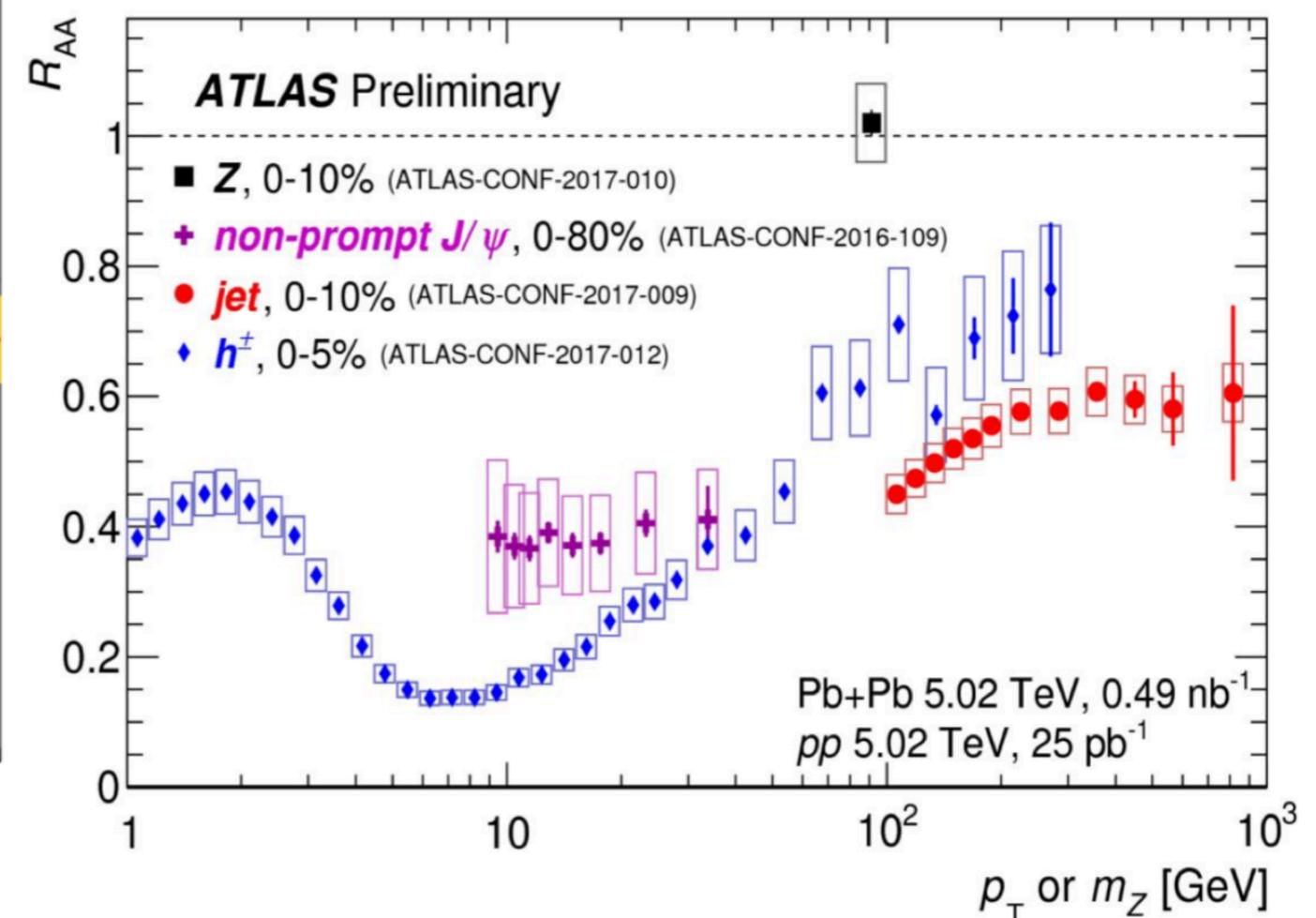
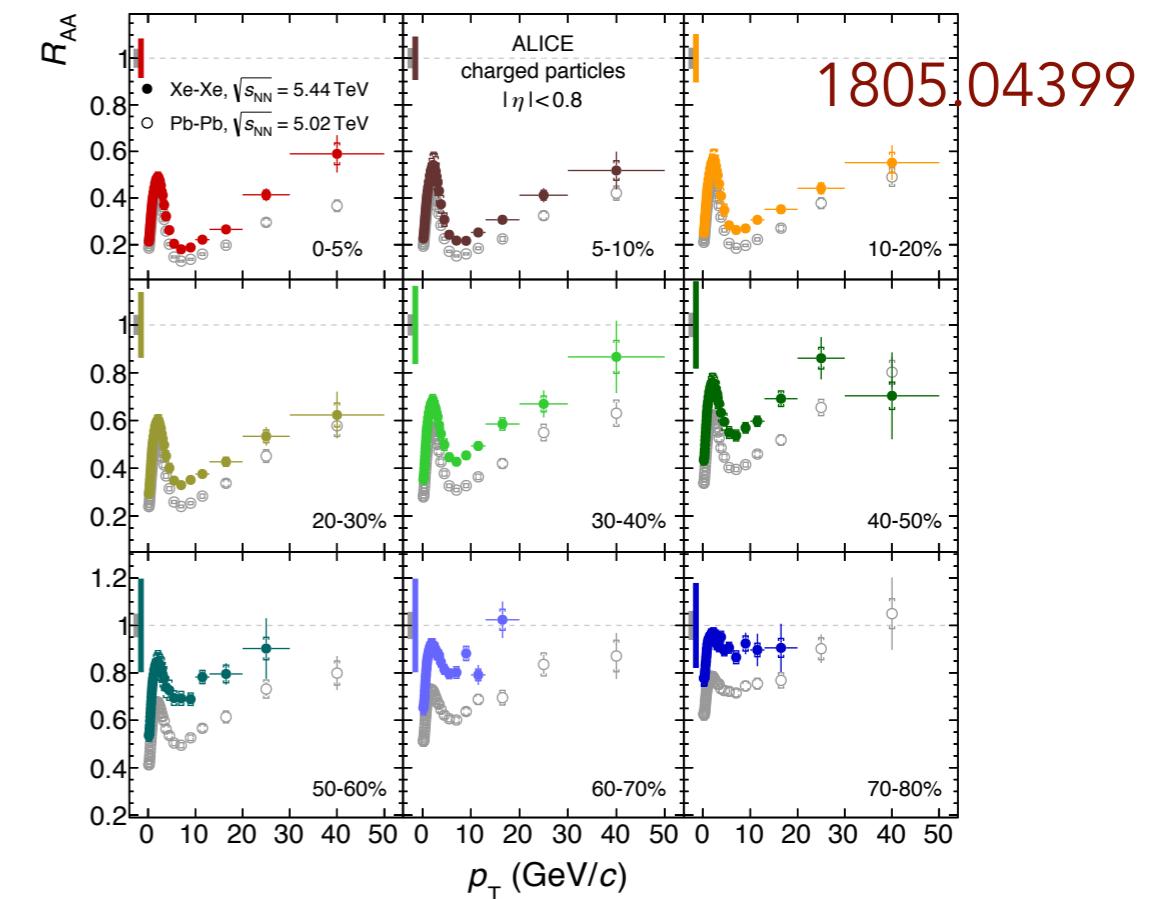
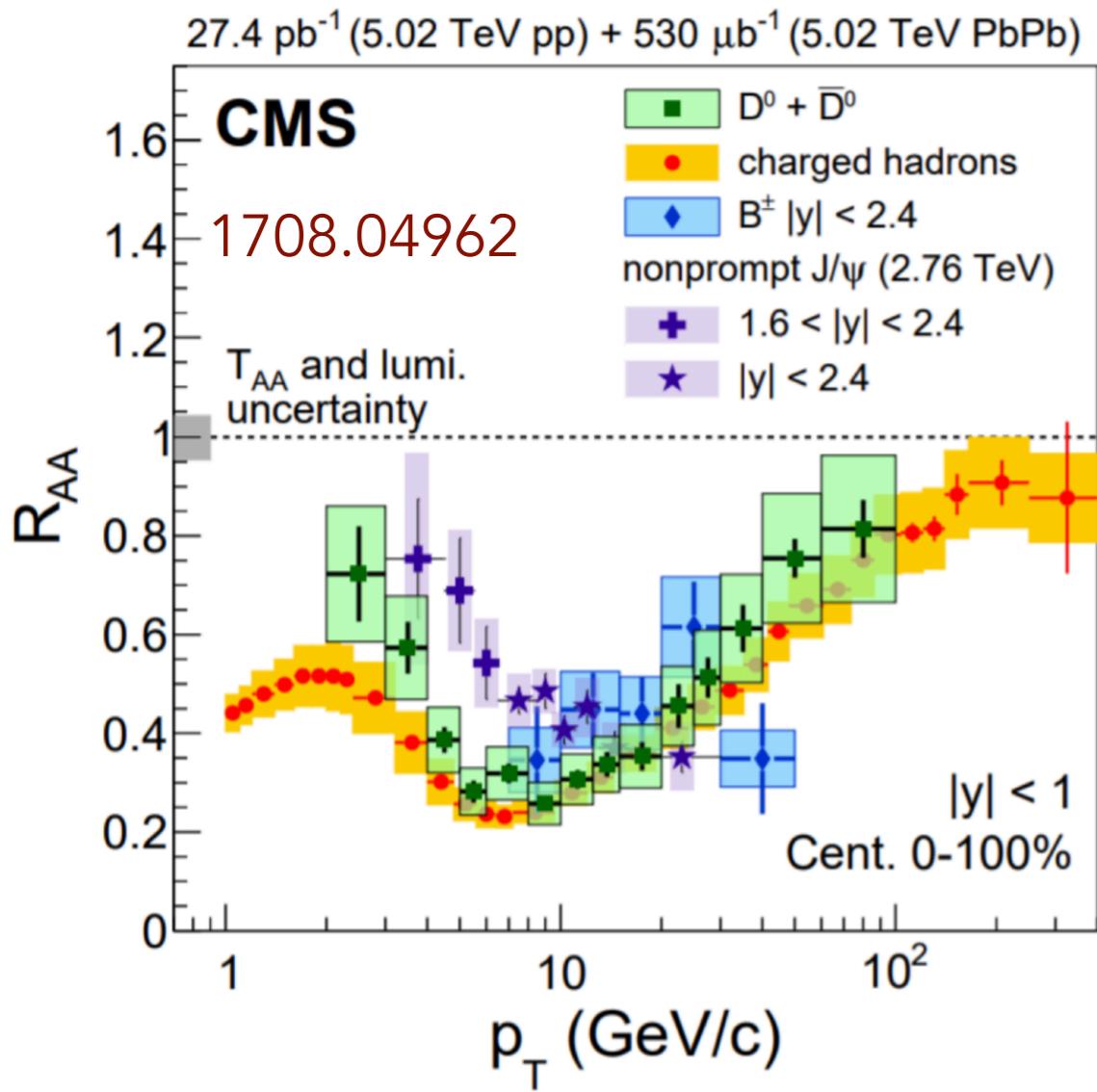
Indian-summer School of Physics,
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3) ENERGY LOSS

- single-inclusive spectrum
- hadrons (quarks)
- heavy quarks
- jets

JET QUENCHING

$$R_{AA} = \frac{dN_{AA}/dp_T^2 dy}{\langle T_{AA} \rangle d\sigma_{pp}/dp_T^2 dy}$$



JET QUENCHING

- strong suppression of wide range of probes of large range in p_T
- sensitive to color charge
- interesting centrality dependence
- common features/unified description?
 - charged hadrons (light quarks/gluons)
 - heavy mesons (heavy quarks/gluons)
 - jets

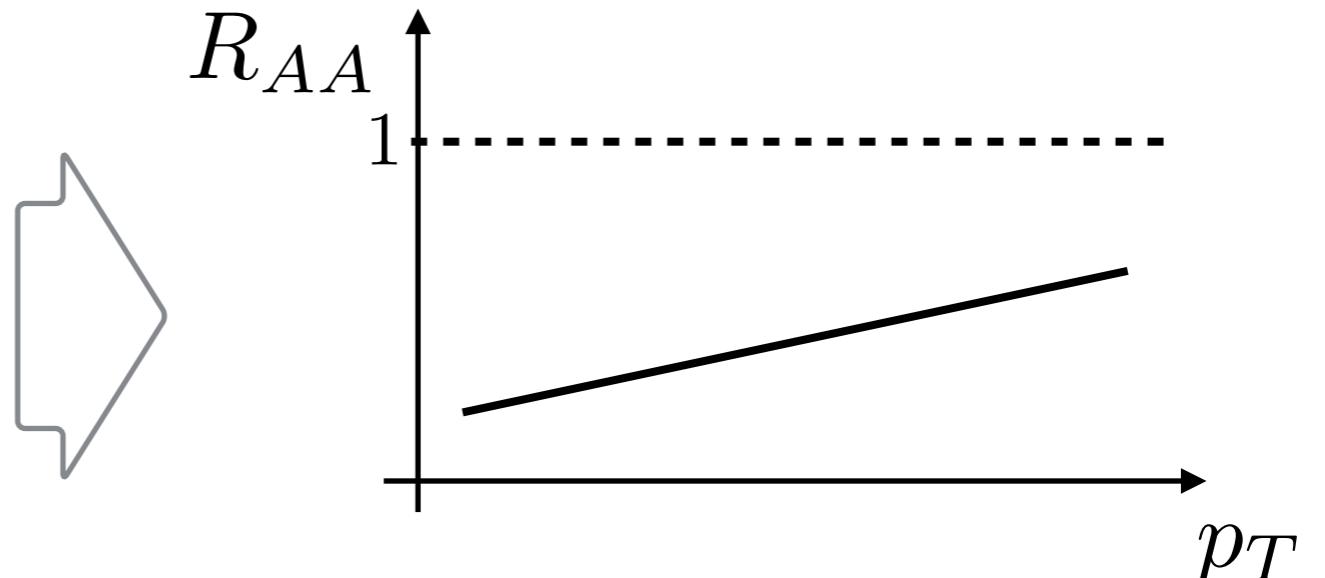
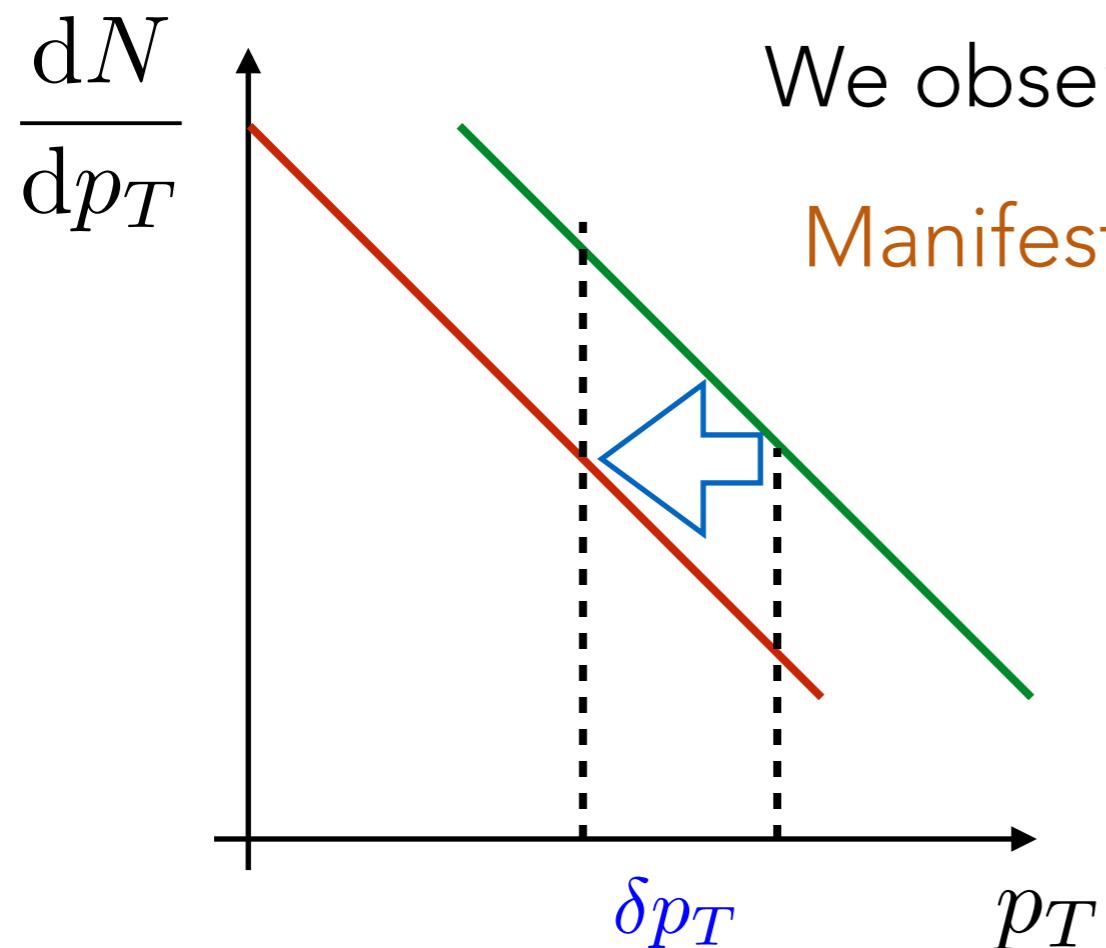
The goal of studying jets is to extract information about the properties of the medium; independent and complementary probes!

WHAT IS GOING ON?

Imagine we shoot a well-controlled flux of particles into the plasma.

We observe suppression because of pT shift.

Manifestation of a energy loss mechanism!



$$\frac{dN_{AA}(p_T)}{dp_T} = \frac{dN_{pp}(p'_T = p_T + \delta p_T)}{dp'_T} \times \left| \frac{dp'_T}{dp_T} \right|$$

GENERALITIES

Our best candidate:
radiative energy loss!

$$\omega \frac{dI}{d\omega} = \frac{\alpha_s C_R}{\pi} \sqrt{\frac{\hat{q}L^2}{\omega}}$$

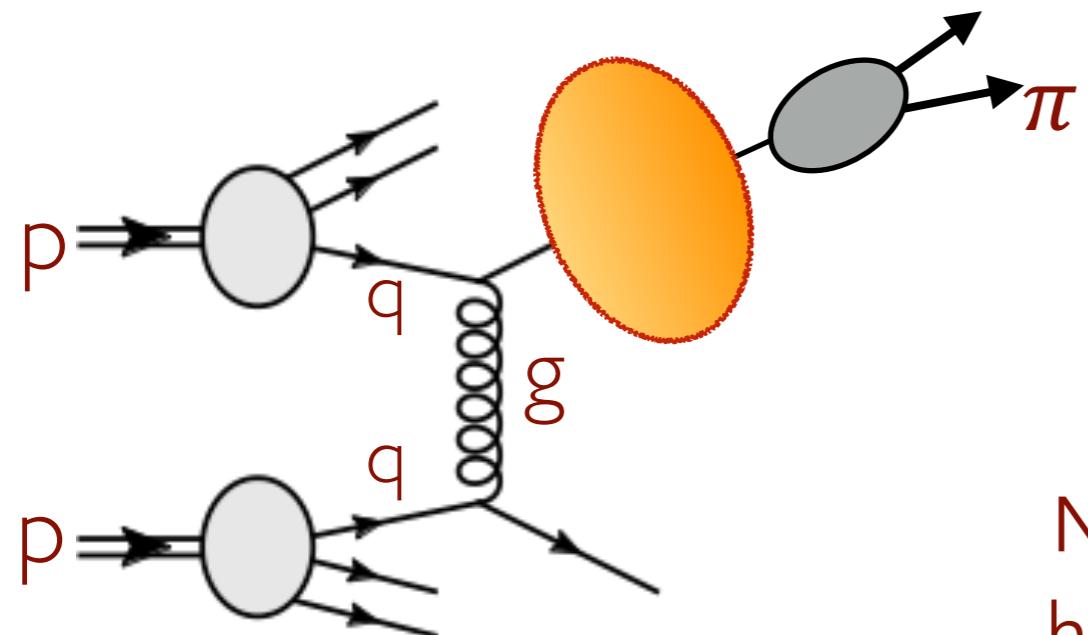
- any color object propagating through the medium sources radiation, leads to energy loss
- but, in principle, other processes can also contribute [elastic energy loss, broadening & absorption]

Probability for losing energy
from one emission

$$\mathcal{P}(\epsilon) = \delta(\epsilon) \left(1 - \int d\omega \frac{dI}{d\omega} \right) + \frac{dI}{d\epsilon}$$

Problem gets more complicated if we consider more sources of radiation
(vacuum radiation) and multi-gluon medium-induced radiation!

QUENCHING OF HADRONS I



Assume single quark/gluon propagates through the medium and fragments outside!

Not a bad approximation since high- p_T hadrons come from hard fragmentation (hadrons \sim narrow jets).

$$\frac{d\sigma_{\text{med}}^h}{dp_T^2 dy} = \int dq_T^2 dz \frac{d\sigma_{\text{vac}}^k}{dq_T^2 dy} D^{k \rightarrow h}(z, \mu_F^2) \mathcal{P}_k(\epsilon) \delta(p_T - z(q_T - \epsilon))$$

$$q_T = p_T/z + \epsilon \quad \text{steeply falling spectrum... hence } z \sim 1$$

Refinement: can also replace $\langle z \rangle$ in the delta function, and integrate out z .

QUENCHING OF HADRONS II

quenching weight: probability distribution of losing energy

$$\frac{d\sigma_{\text{med}}}{dp_T^2 dy} = \int_0^\infty d\epsilon \mathcal{P}(\epsilon) \frac{d\sigma_{\text{vac}}(p_T + \epsilon)}{dp_T^2 dy}$$

quenching factor = nuclear modification factor

$$R_{\text{jet}} = \left(\frac{d\sigma_{\text{med}}}{dp_T^2 dy} \right) / \left(\frac{d\sigma_{\text{vac}}}{dp_T^2 dy} \right)$$

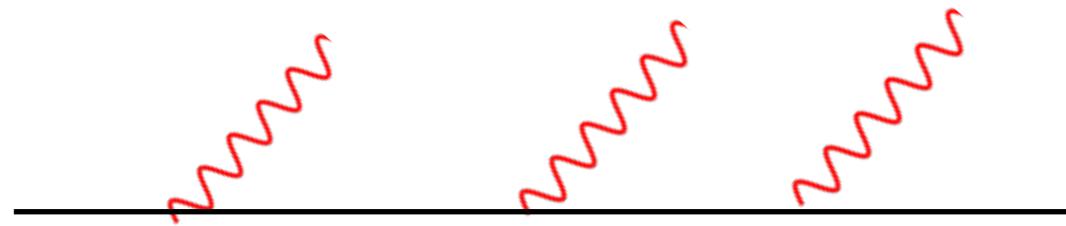
For $\epsilon/p_T \ll 1$ and large n : $\frac{1}{(p_T + \epsilon)^n} \approx \frac{1}{p_T^n} e^{-\epsilon n/p_T}$

$$R_{\text{jet}} \sim \tilde{\mathcal{P}}(n/p_T) \equiv \mathcal{Q}(p_T)$$

quenching factor is Laplace transform of energy loss probability

[For now, $n=\text{const}$, but this can be improved and finite- n corrections can be implemented.]

RESUMMATION OF MULTIPLE GLUONS



Soft gluons: copious, large angle, short formation time

no interferences!

$$\mathcal{P}(\epsilon) = \delta(\epsilon) \left(1 - \int d\omega \frac{dI}{d\omega} \right) + \frac{dI}{d\epsilon}$$



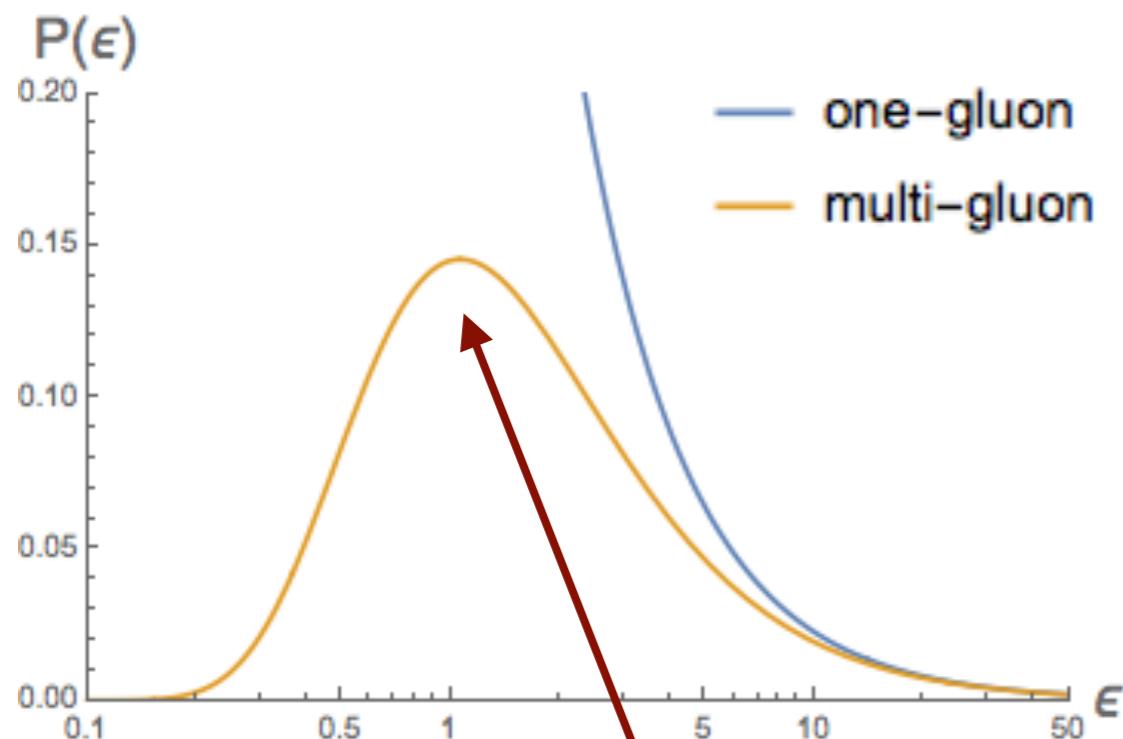
$$\mathcal{P}(\epsilon, L) = \delta(\epsilon) + \int_0^L dt \int d\omega \left[\frac{dI}{d\omega dt} - \delta(\omega) \int d\omega' \frac{dI}{d\omega' dt} \right] \mathcal{P}(\epsilon - \omega, t)$$

$$\frac{dI}{d\omega dt} \underset{\sim}{=} \frac{\bar{\alpha}}{\omega t_{\text{br}}(\omega)}$$

Constant rate for infinite medium;
time-dependent rate in general

RATE EQUATION

$$\frac{\partial}{\partial t} \mathcal{P}(\epsilon, t) = \int d\omega \left[\frac{dI}{d\omega dt} - \delta(\omega) \int d\omega' \frac{dI}{d\omega' dt} \right] \mathcal{P}(\epsilon - \omega, t)$$



E-loss dominated by max of distribution!

$$\tilde{f}(\nu) = \int_0^\infty d\omega e^{-\nu\omega} f(\omega)$$

$$\frac{\partial}{\partial t} \tilde{\mathcal{P}}(\nu, t) = \gamma(\nu, t) \tilde{\mathcal{P}}(\nu, t)$$

$$\begin{aligned} \gamma(\nu, t) &= \int_0^\infty d\omega (e^{-\nu\omega} - 1) \frac{dI}{d\omega dt} \\ &= -2\sqrt{\pi\nu \bar{\alpha}^2 \hat{q}} \end{aligned}$$

$$\mathcal{P}(\epsilon) = \sqrt{\frac{\omega_s}{\epsilon^3}} e^{-\frac{\pi\omega_s}{\epsilon}}$$

QUARK QUENCHING FACTOR

$$\mathcal{Q}(p_T) = e^{-2\bar{\alpha}L\sqrt{n\pi\hat{q}/p_T}} \quad \bar{\alpha} = \frac{\alpha_s C_R}{\pi}$$

$$\mathcal{Q}_g(p_T) = (\mathcal{Q}_q(p_T))^{N_c/C_F}$$

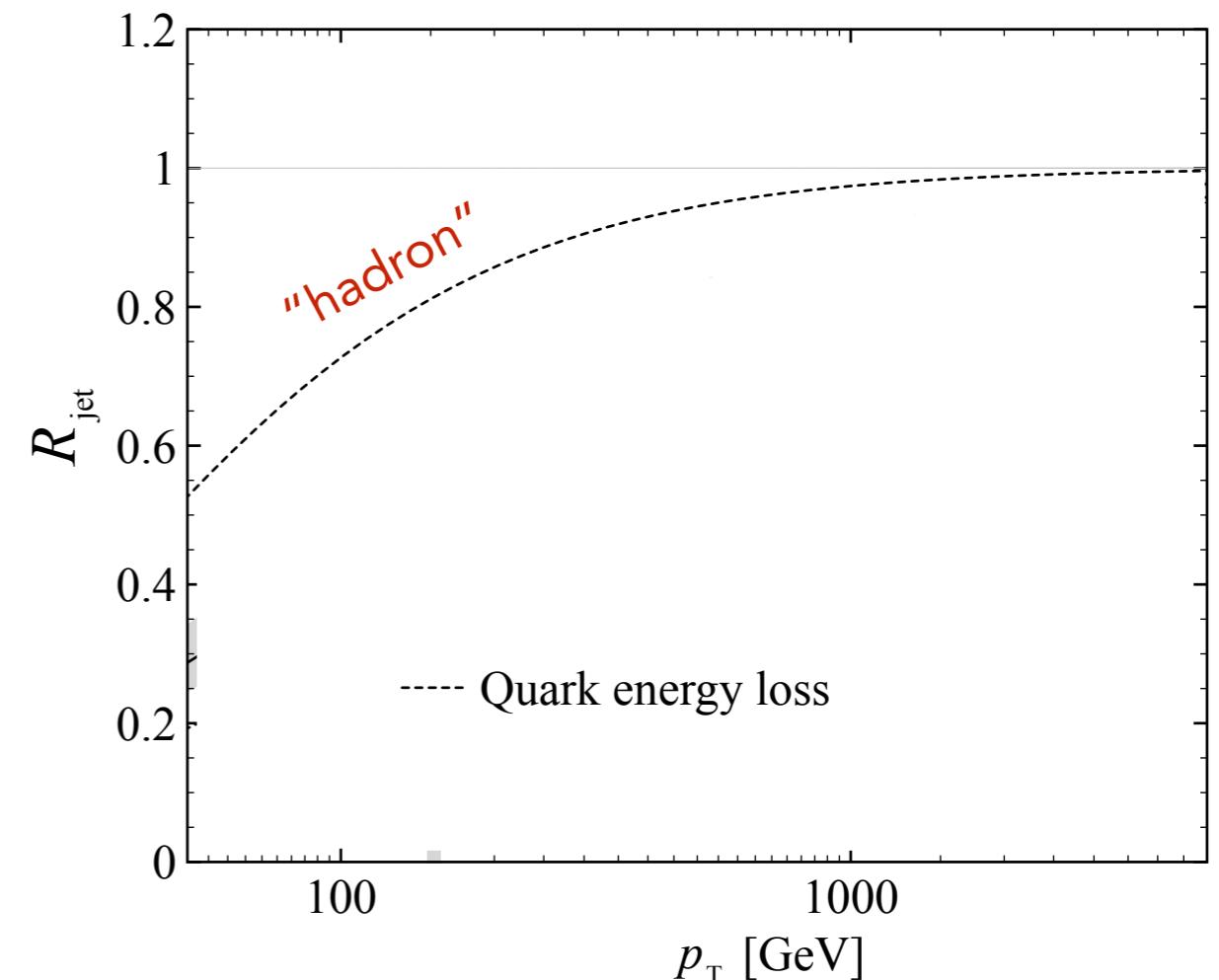
- form factor suppression related to multiplicity of virtual gluons

$$\mathcal{Q}(p_T) \sim e^{-N(\omega>p_T/n)}$$

- strong quenching for

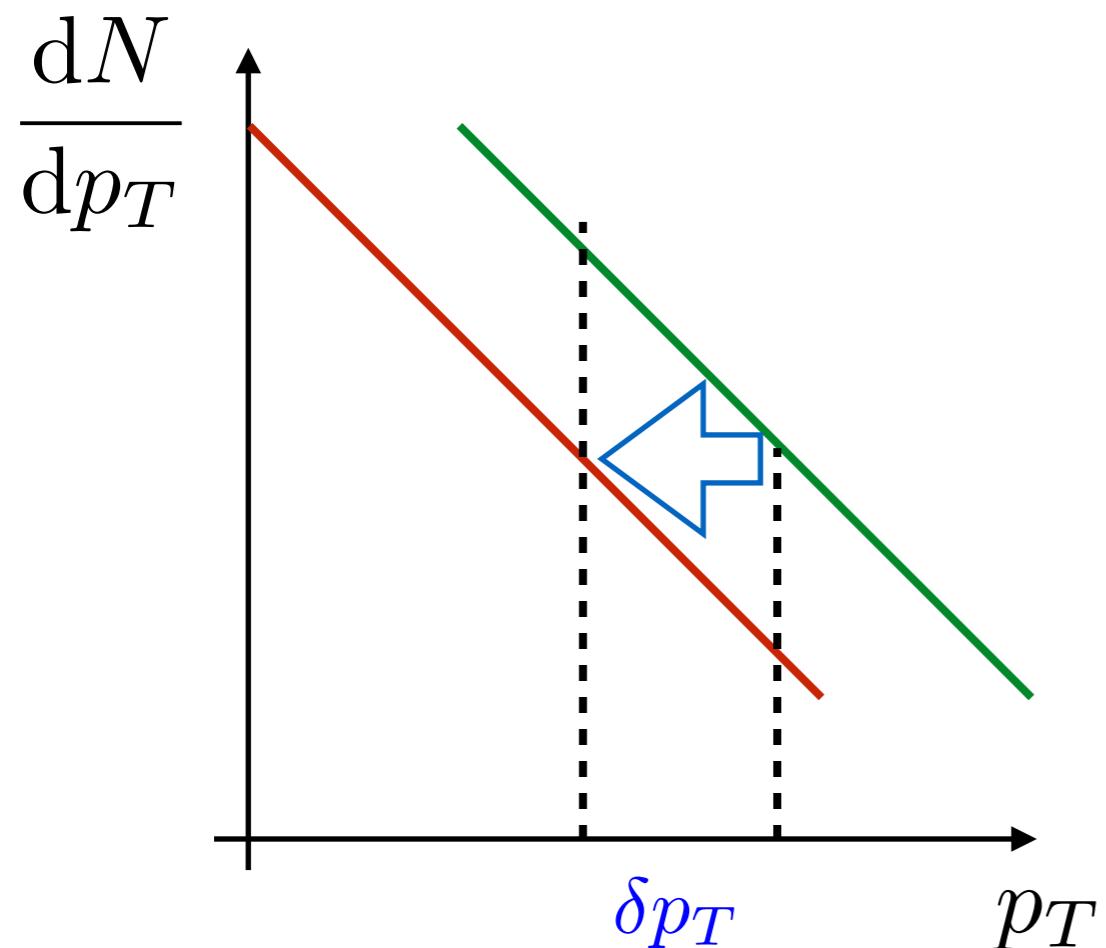
$$p_T \ll n\bar{\alpha}^2\hat{q}L^2$$

- scale related to multiplicity of emitted gluons (not energy)



SHIFTING THE SPECTRUM

Our analysis have resulted in a "model" w/



$$\delta p_T = \sqrt{\frac{8\pi\bar{\alpha}^2\hat{q}L^2 p_T}{n}}$$

- for a static medium,....

Recall, that naïve expectation

$$\delta p_T = \bar{\alpha}\hat{q}L^2$$

This only holds in the high-pT limit!

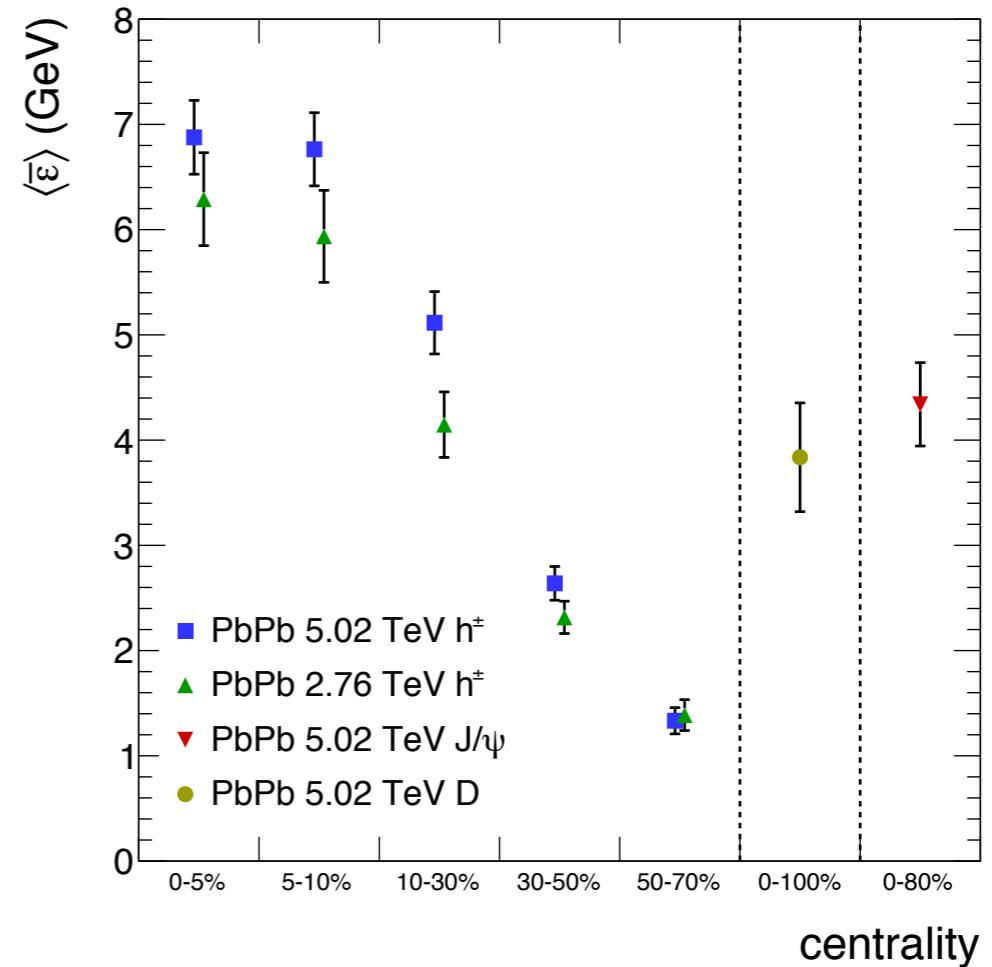
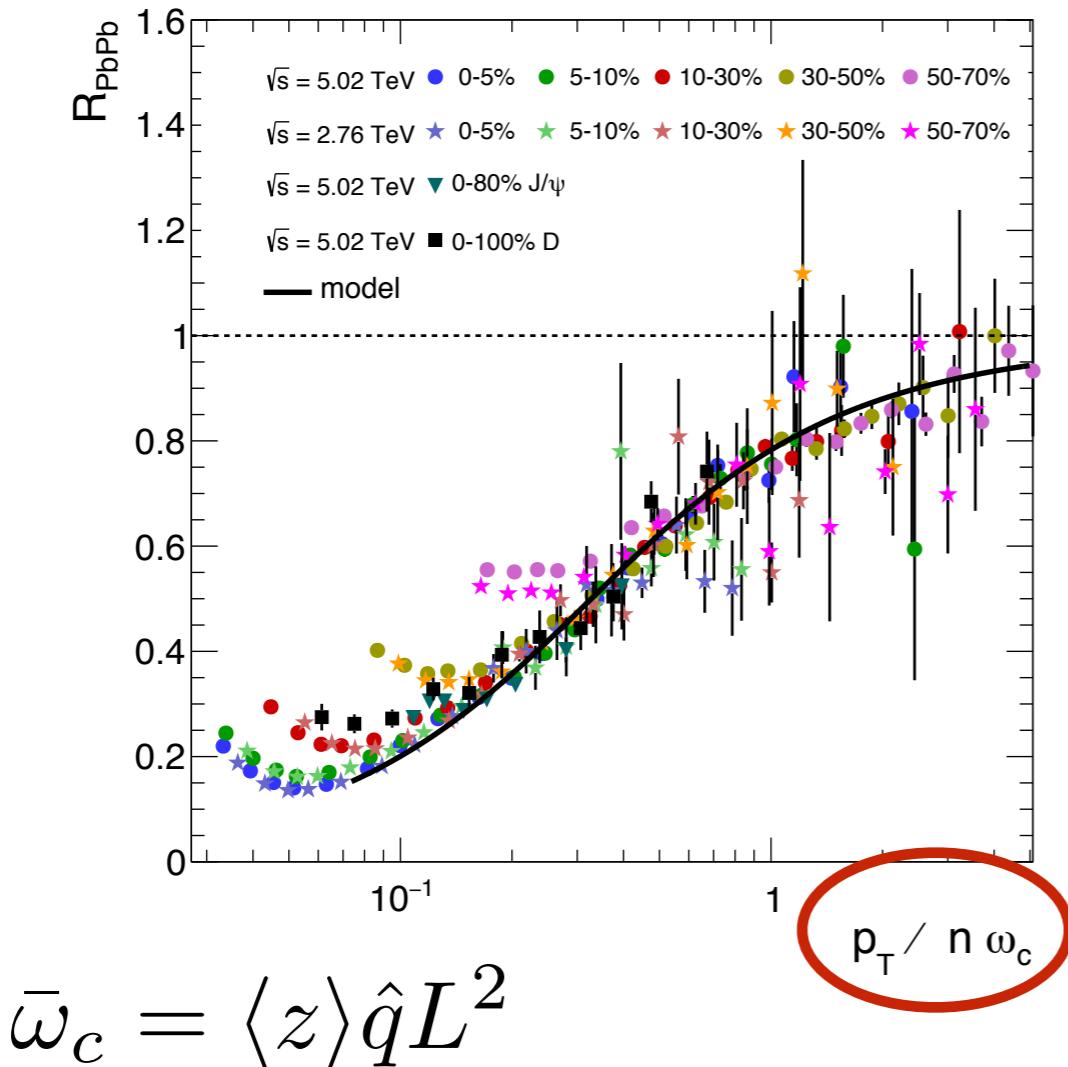
$$p_T > n\hat{q}L^2$$

But in that limit, quenching is weak because the **multiplicity** is small!

$$\mathcal{Q}(p_T > n\hat{q}L^2) \sim \exp[-N(\omega > \omega_c)] \sim e^{-\bar{\alpha}} \sim 1 - \bar{\alpha}$$

SCALING FEATURE

Arleo (2017)

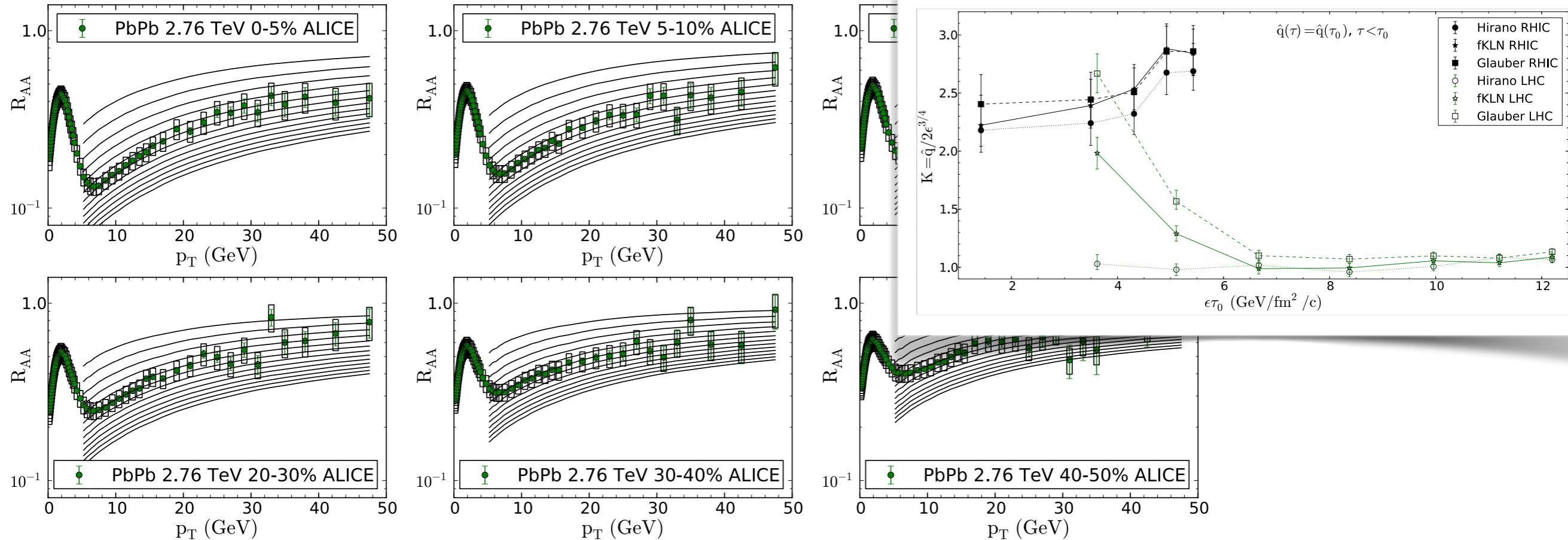


Due to the convolution with a steeply falling spectrum, we are only sensitive to the **combination** - valid even for **heavy particles**!

Breaks down at low p_T (small n)...

PHENOMENOLOGICAL ANALYSIS

Andres, Armesto, Luzum, Salgado, Zurita (2016)



Extracting \hat{q} from data (w/ realistic hydro, FF, NLO etc..) $\hat{q}(\xi) = K \cdot 2\epsilon^{3/4}(\xi)$

State of the art calculations allow to extract medium properties, check for consistency...

HEAVY-QUARK QUENCHING

$$dN^{\text{vac}} = \bar{\alpha} \frac{dz}{z} \frac{k_\perp^2 dk_\perp^2}{[k_\perp^2 + (zm)^2]^2} = \bar{\alpha} \frac{d\omega}{\omega} \frac{d\theta^2}{\theta^2} \left(1 + \frac{\theta_0^2}{\theta^2}\right)^{-2}$$



"dead cone" effect
 $\theta_0 = m/p_T$

$$dN^{\text{med}} = \bar{\alpha} \frac{d\omega}{\omega} \sqrt{\frac{\hat{q}L^2}{\omega}} \left(1 + \frac{\theta_0^2}{\theta_{\text{br}}^2(\omega)}\right)^{-2}$$

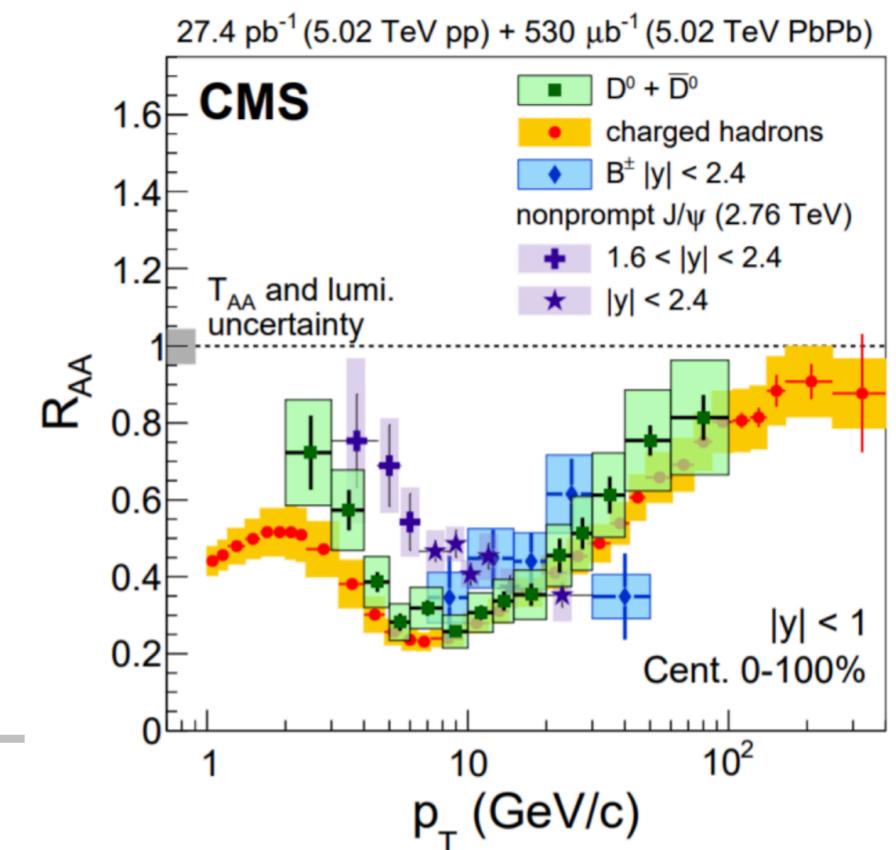
Dokshitzer, Kharzeev (2001); Armesto, Salgado Wiedeman (2003)

0th order expectation:

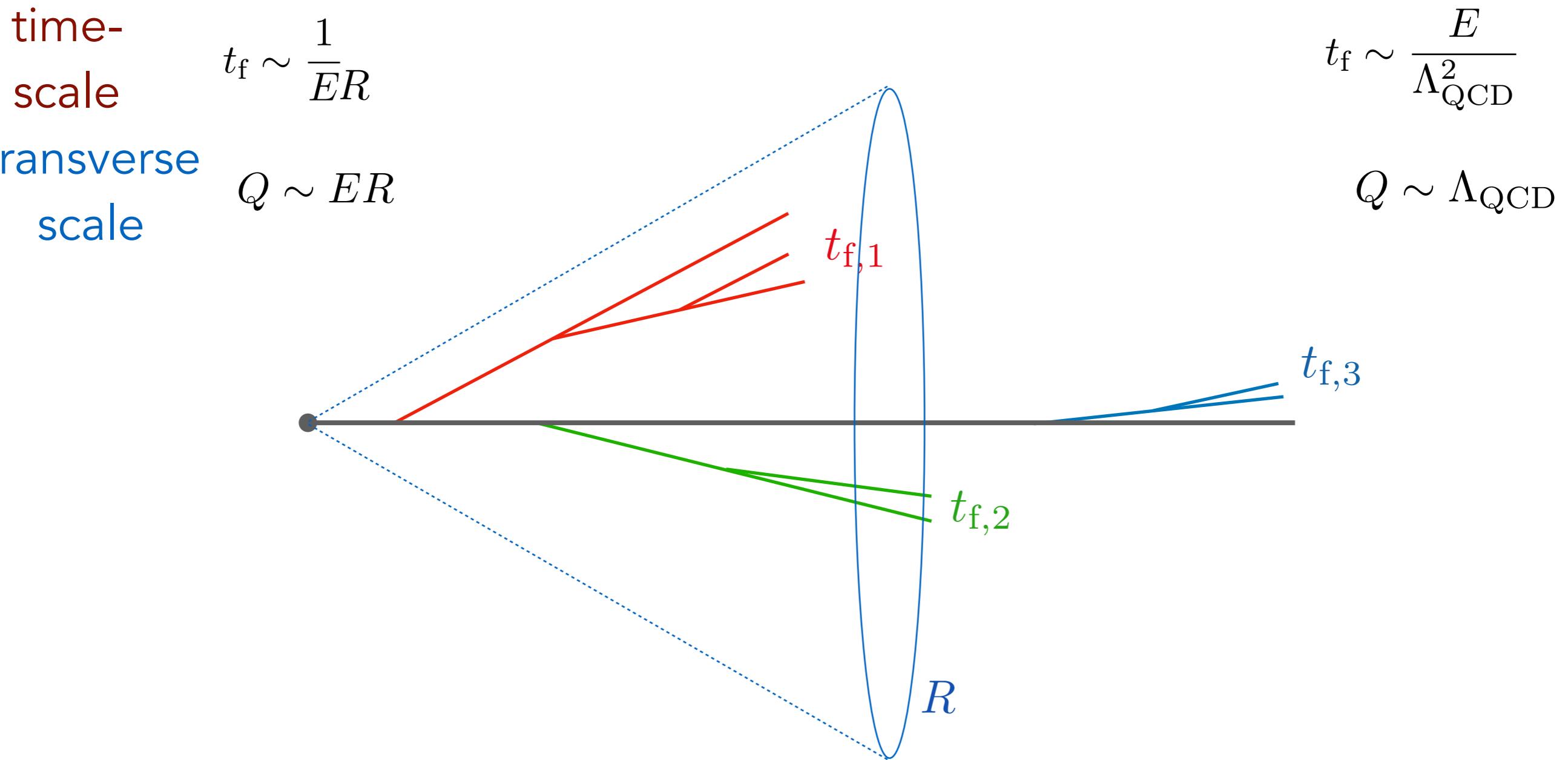
$$\frac{\mathcal{Q}(p_T, m)}{\mathcal{Q}(p_T, 0)} \simeq \exp \left[2\bar{\alpha} L \sqrt{\pi} \left(\frac{\hat{q}}{\theta_0^2} \right)^{1/3} \right]$$

Heavy-quarks should be less quenched than light!

Data: realistic spectrum, effect of elastic e-loss



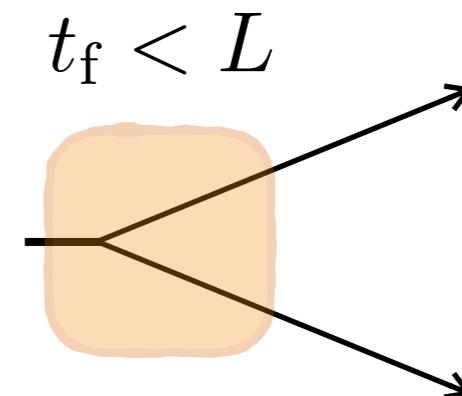
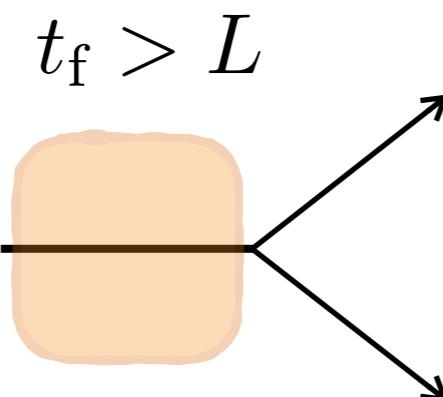
WHAT ABOUT JETS?



Generally have to deal with multi-gluon emissions!

Q: did the gluons form inside or outside of the medium?

A SIMPLE MODEL



Quenching effect very different in the two situations...
(quenching of two vs. quenching of one)

Let's assume that the medium quenches completely all
emissions inside the medium!

$$N(t_f < L) = \bar{\alpha} \int_{(\omega\theta^2)^{-1} < L} \frac{d\omega}{\omega} \frac{d\theta}{\theta} = \frac{\bar{\alpha}}{2} \ln^2 p_T R L$$

could affect many splittings!

RADIATIVE CORRECTIONS

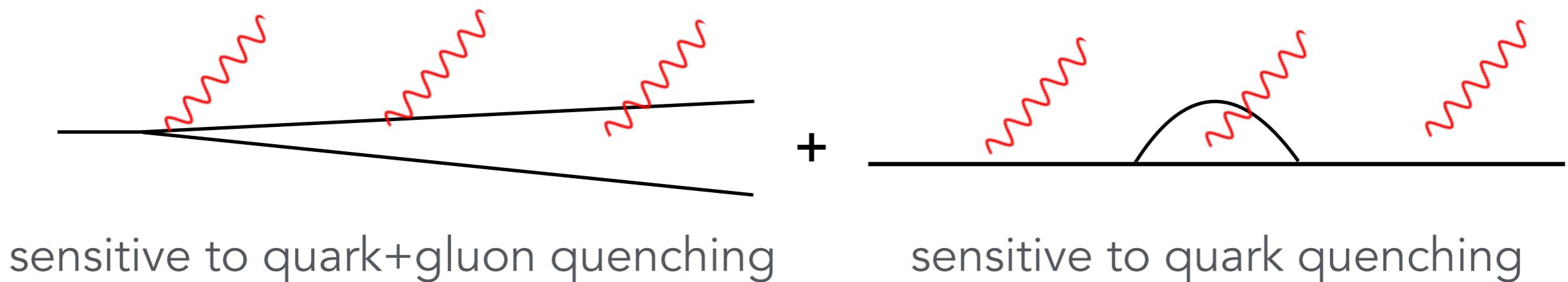
$$\frac{d\sigma}{dydp_T^2} = \frac{d\sigma_{\text{Born}}}{dydp_T^2} \left[1 + \alpha_s \left(\int d\Pi_{\text{real}} - \int d\Pi_{\text{virt}} \right) + \mathcal{O}(\alpha_s^2) \right]$$

- higher-order corrections not enhanced by phase space when balance between **real** & **virtual** emissions for sufficiently **inclusive observables**
- does this hold in the medium (final-state interactions)?

$$R_{\text{jet}} = Q^{(0)}(p_T) + Q^{(1)}(p_T) + \mathcal{O}(\alpha_s^2)$$

- expanding quenching factor corresponds to accounting for the quenching of **higher-order vacuum emissions** (substructure fluctuations)

HIGHER-ORDER CORRECTION TO QUENCHING



$$\begin{aligned} Q^{(1)}(p_T) &= \int_0^1 dz P_{gq}(z) \int_0^R \frac{d\theta}{\theta} \frac{\alpha_s(k_\perp)}{\pi} [Q_{gq}(p_T) - Q_q(p_T)] \\ &\simeq -Q_q(p_T) N(t_f < L) \end{aligned}$$

- real & virtual are differently affected by energy loss effects!
- the mismatch is largest at short formation times

Our model simply kills all real emissions!

SUDAKOV SUPPRESSION OF JETS

$$R_{\text{jet}} = \mathcal{Q}_q(p_T) \times \mathcal{C}(p_T, R)$$

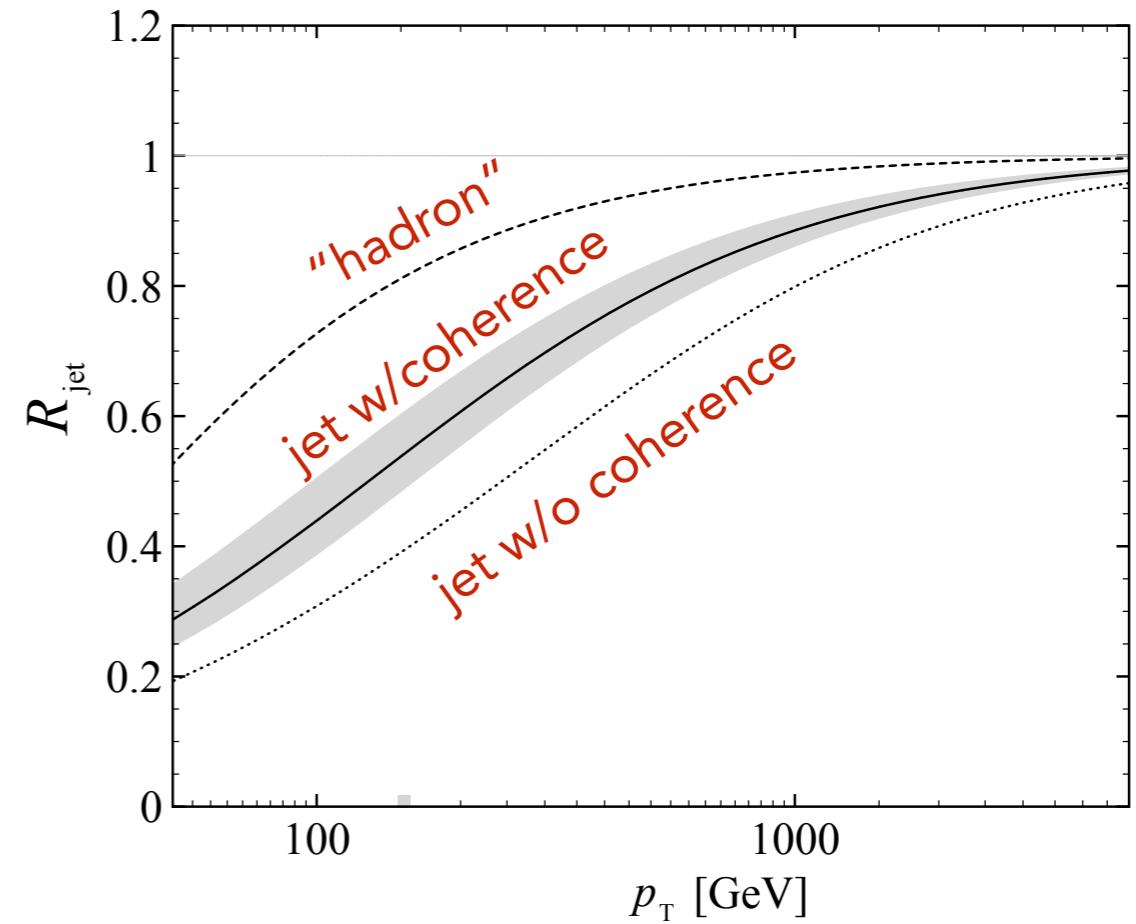
jet loses energy via **total charge** & **resolved substructure fluctuations**

Strong quenching:

Sudakov form factor, depends
explicitly on **jet scales**

- finite quenching: non-linear evolution equation for $\mathcal{C}(p_T, L)$
- minimal resolution angle $\sim \theta_c$ of the medium (coherence)
- hierach sensitive to jet scales

$$\mathcal{C}(p_T, R) \simeq e^{-N(t_f < L)}$$



SUMMARY

- the spectra of particles (light, heavy) and jets are suppressed due to interactions with the medium
 - can be experimentally accessed as a δp_T shift
- radiative energy loss (sensitive to mass)
- jets are multi-gluon objects: additional quenching
- 0th order expectation is a hierarchy

$$R_Q > R_q > R_{\text{jet}}$$

OUTLOOK

- jets are interesting physical objects at colliders
 - calibrated: theoretical tools & MC
- radiative processes in the medium leads to energy loss
- different phenomenological implications for single-inclusive spectra for hadrons & jets
- also, di-jet/boson-jet great for pinning down short- and long-distance pieces
- promising new observables: jet substructure